EXERCISE 18

# Patterns in the Solar System

Although composed of many diverse objects, the solar system (Figure 18.1) exhibits various degrees of order and several regular patterns. To simplify the investigation of planetary sizes, masses, etc., the planets can be arranged into two distinct groups, with the members of each displaying similar attributes. This exercise examines the physical properties and motions of the planets with the goal of summarizing these characteristics in a few general, easily remembered statements.

## **Objectives**

After you have completed this exercise, you should be able to:

- **1.** Describe the appearance of the solar system when it is viewed along the plane of the ecliptic.
- **2.** Summarize the distances and spacing of the planets in the solar system.

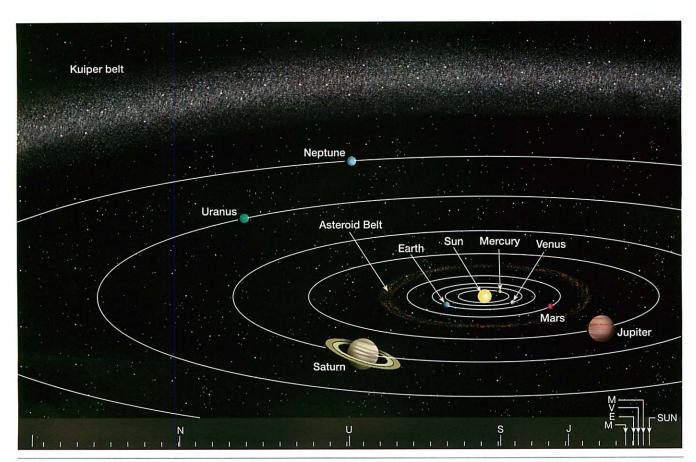


Figure 18.1 The solar system showing the orbits of the planets to scale. A different scale has been used for the sizes of the Sun and planets. Therefore, the diagram is not a true scale model representation of the solar system.

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- **3.** Summarize and compare the physical characteristics of the terrestrial and Jovian planets.
- **4.** Describe the motions of the planets in the solar system.

### **Materials**

ruler colored pencils calculator

Materials Supplied by Your Instructor

4-meter length of adding machine paper with covers and meterstick thermometers light source (150-watt bulb)

### **Terms**

nebula	mass	revolution
terrestrial planets	density	Kepler's laws
Jovian planets	weight	astronomical unit
dwarf planets	rotation	
plane of the ecliptic		

### Introduction

The order that exists within the solar system is directly related to the laws of physics that governed its formation. Astronomers have determined that the Sun and planets originated approximately 4.6 billion years ago from an enormous cloud of dust and gas. As this **nebula** contracted, it began to rotate and flatten. Eventually the temperature and pressure in the center of the cloud was great enough to initiate nuclear fusion and form the Sun.

Near the center of the nebula, the planets Mercury, Venus, Earth, and Mars evolved under nearly the same conditions and consequently exhibit similar physical properties. Because these planets are rocky objects with solid surfaces, they are collectively called the terrestrial (Earth-like) planets.

The outer planets, Jupiter, Saturn, Uranus, and Neptune, being farther from the Sun than the terrestrial planets, formed under much colder conditions and are gaseous objects with central cores of ices and rock. Since the four planets are very similar, they are often grouped together and called the **Jovian (Jupiterlike)** planets.

The **dwarf planets**, which include, among other celestial objects, Pluto and the former asteroid Ceres, are not included with either the terrestrial or Jovian planets.

Table 18.1 illustrates many of the individual characteristics of the planets in the solar system.

- 1. Examine the data in Table 18.1, then
  - a. Draw lines on the upper and lower parts of Table 18.1 that separate the terrestrial planets from the Jovian planets. Label the lines "Belt of Asteroids."

Table 18.1 Planetary data.

		Mea	an Distance from	Sun				oital ocity
Planet	Symbol	AU	Millions of Miles	Millions of Kilometers	Period of Revolution	Inclination of Orbit	mi/s	km/s
Mercury	Ď	0.387	36	58	88 <sup>d</sup>	7°00′	29.5	47.5
Venus	9	0.723	67	108	224.7 <sup>d</sup>	3°24′	21.8	35.0
Earth	$\oplus$	1.000	93	150	365.25 <sup>d</sup>	0°00′	18.5	29.8
Mars	ð	1.524	142	228	687 <sup>d</sup>	1°51′	14.9	24.1
Jupiter	24	5.203	483	778	11.86 <sup>yr</sup>	1°18′	8.1	13.1
Saturn	ħ	9.539	886	1427	29.46 <sup>yr</sup>	2°29′	6.0	9.6
Uranus	8	19.180	1783	2870	84 <sup>yr</sup>	0°46′	4.2	6.8
Neptune	Ψ	30.060	2794	4497	165 <sup>yr</sup>	1°46′	3.3	5.3

	Period of	Diameter		Relative - Mass	Average Density	Polar	Mean	Number of
Planet	Rotation	Miles	Kilometers	(Earth = 1)	(g/cm <sup>3</sup> )	Flattening (%)	Temperature (°C)	Known Satellites
Mercury	59 <sup>d</sup>	3015	4854	0.056	5.4	0.0	167	0
Venus	244 <sup>d</sup>	7526	12,112	0.82	5.2	0.0	464	0
Earth	23 <sup>h</sup> 56 <sup>m</sup> 04 <sup>s</sup>	7920	12,751	1.00	5.5	0.3	15	1
Mars	24h37m23s	4216	6788	0.108	3.9	0.5	-65	2
Jupiter	9 <sup>h</sup> 50 <sup>m</sup>	88,700	143,000	317.87	1.3	6.7	-110	63
Saturn	10 <sup>h</sup> 14 <sup>m</sup>	75,000	121,000	95.14	0.7	10.4	-140	56
Uranus	17 <sup>h</sup> 14 <sup>m</sup>	29,000	47,000	14.56	1.2	2.3	-195	27
Neptune	16 <sup>h</sup> 03 <sup>m</sup>	28,900	46,529	17.21	1.7	1.8	-200	13

b. On both parts of the table write the word "terrestrial" next to Mercury, Venus, Earth, and Mars and the word "Jovian" next to Jupiter, Saturn, Uranus, and Neptune.

# The "Shape" of the Solar System

When the solar system is viewed from the side, the orbits of the planets all lie in nearly the same plane, called the plane of the ecliptic (Figure 18.1). The column labeled "Inclination of Orbit" in Table 18.1 lists how many degrees the orbit of each planet is inclined from the plane. Answer questions 2-4 by referring to Table 18.1.

- 2. Other than Mercury, whose orbit is inclined (4, 7, 10) degrees, the orbits of the remaining planets are all within (4, 7, 10) degrees of the plane of the ecliptic. Circle your answers.
- 3. The orbit of the dwarf planet Pluto is inclined 17° to the plane. When compared to the eight planets, Pluto's orbit is:
- 4. Considering the nebular origin of the solar system, suggest a reason why the orbits of the planets are nearly all in the same plane.

# Distance and Spacing of the Planets

An examination of any scale-model solar system reveals that the distances from the Sun and the spacing between the planets appear to follow a regular pattern. Although many ancient astronomers were concerned with planetary distances and spacing, it was not until the mid-1700s that astronomers found a simple mathematical relation that described the arrangement of the planets known at the time.

#### A Scale Model of Planetary Distances

Perhaps the best way to examine distance and spacing of the planets in the solar system is to use a scale model.

- 5. Prepare a distance scale model of the solar system according to the following steps.
- Step 1: Obtain a 4-meter length of adding machine paper and a meterstick from your instructor.
- Step 2: Draw an "X" about 10 centimeters from one end of the adding machine paper and label it "Sun."

Step 3: Using the mean distances of the planets from the Sun in miles presented in Table 18.1 and the following scale, draw a small circle for each planet at its proper scale mile distance from the Sun. Use a different colored pencil for the terrestrial and Jovian planets and write the name of the planet next to its position.

#### **SCALE**

1 millimeter = 1 million miles 1 centimeter = 10 million miles 1 meter = 1,000 million miles

Step 4: Write the word "asteroids" 258 million scale miles from the Sun.

Answer questions 6-9 using the distance scale model you constructed in question 5.

- 6. What feature of the solar system separates the terrestrial planets from the Jovian planets?
- 7. Observe the scale model diagram and summarize the spacing for each of the two groups of planets.

Spacing of the terrestrial planets:	
Spacing of the Jovian planets:	

- 8. Write a brief statement that describes the spacing of the planets in the solar system.
- 9. Which planet(s) vary the most from the general pattern of spacing?

## **Comparing the Terrestrial** and Jovian Planets

The physical characteristics such as diameter, density, and mass of the terrestrial planets are very similar and can be summarized in a few statements. Likewise, the characteristics exhibited by the Jovian planets as a group can also be generalized.

To gain an understanding of the similarities of the planets within each of the two groups and the contrasts between the two groups, complete the following sections using the planetary data presented in Table 18.1.

#### Size of the Planets

The similarities in the diameters of the planets within each of the two groups and the contrast between the groups are perhaps the most obvious patterns in the solar system. The diameter of each planet is given in both miles and kilometers in Table 18.1.

- 10. To visually compare the relative sizes of the planets and Sun, complete the following steps using the unmarked side of your 4-meter length of adding machine paper.
- **Step 1:** Determine the radius of each planet in kilometers by dividing its diameter (in kilometers) by 2. List your answers in the "Radius" column of Table 18.2.
- Step 2: Use a scale of 1 cm = 2,000 km. Determine the scale model radius of each planet and list your answer in the "Scale Model Radius" column of Table 18.2.
- **Step 3:** Draw an "X" about 10 cm from one end of the adding machine paper and label it "Starting point."
- Step 4: Using the scale model radius in Table 18.2, begin at the starting point and mark the radius of each planet with a line on the paper. Use a different colored pencil for the terrestrial and Jovian planets. Label each line with the planet's name.
- Step 5: The diameter of the Sun is approximately 1,350,000 kilometers. Using the same scale as you used for the planets (1 cm = 2000 km), determine the scale model radius of the Sun. Mark the Sun's radius on the adding machine paper using a different colored pencil from the two planet groups. Label the line "Sun."

Table 18.2 Planetary Radii with Scale Model Equivalents

Planet	Radius (in kilometers)	Scale Model Radius
Mercury		cm
Venus		cm
Earth		cm
Mars		cm
Jupiter		cm
Saturn		cm
Uranus		cm
Neptune		cm

Answer questions 11–17 using both Table 18.1 and the scale model radius diagram you constructed in question 10.

11.	Which is the largest of the terrestrial planets and what is its diameter?
	miles
12.	Which is the smallest Jovian planet and what is its diameter?
	miles
13.	Complete the following statement.
	The smallest Jovian planet,, is, is times larger than the largest terrestrial planet.
14.	Summarize the sizes of the planets within each group.
	The diameters of the terrestrial planets:
	The diameters of the Jovian planets:
15.	Write a general statement that compares the sizes of the terrestrial planets to those of the Jovian planets.
16.	Complete the following statement.
	The Sun is times larger than Earth and times larger than Jupiter.
17.	The diameter of the dwarf planet Pluto is approximately 1,500 miles, which is about (one-

### Mass and Density of the Planets

Mass is a measure of the quantity of matter an object contains. In Table 18.1 the masses of the planets are given in relation to the mass of Earth. For example, the mass of Mercury is given as 0.056, which means that it consists of only a small fraction of the quantity of matter that Earth contains. On the other hand, the Jovian planets all contain several times more matter than Earth.

smallest planet. Circle your answer.

third, one-half, twice) the diameter of the

**Density** is the mass per unit volume of a substance. In Table 18.1 the average densities of the planets are expressed in grams per cubic centimeter (g/cm<sup>3</sup>). As a reference, the density of water is approximately one gram per cubic centimeter.

Table 18.1, answer questions 18–22. 18. Complete the following statements: a. The planet \_\_\_\_\_\_ is the most massive planet in the solar system. It is \_\_\_\_\_ times more massive than Earth. b. The least massive planet is \_\_ which contains only \_\_\_\_\_ as much mass as Earth. The gravitational attraction of a planet is directly related to its mass. 19. Which planet exerts the greatest, and which the least, pull of gravity? Explain your answer. Your weight is a function of the gravitational attraction of an object on your mass. 20. The surface gravities of Mars and Jupiter are respectively about 0.4 and 2.5 that of Earth. What would be your approximate weight on each of these planets? 21. Which of the two groups of planets would have the greatest ability to hold large quantities of gas as part of their compositions? Explain your answer. **22.** Write a general statement comparing the masses and gravitational attractions of the terrestrial planets to those of the Jovian planets. **Diameter versus Density** 

Using the relative masses of the planets given in

To visually compare the diameters and densities of the planets, use the data in Table 18.1 to complete the diameter versus density graph, Figure 18.2, according to the procedure in question 23.

23. Plot a point on the diameter verses density graph, Figure 18.2, for each planet where its diameter intersects its density. Label each point with the planet's name. Use a different colored pencil for the terrestrial and Jovian planets.

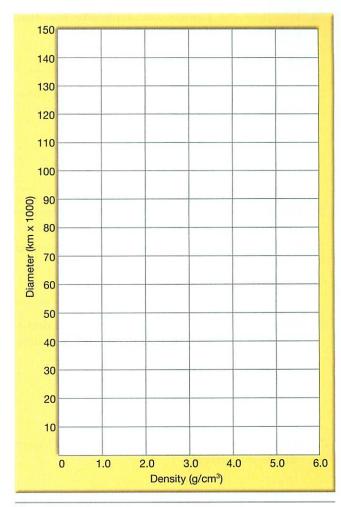


Figure 18.2 Diameter verses density graph.

Answer questions 24–34 using Table 18.1 and the diameter verses density graph you constructed in question 23.

- **24.** What general relation exists between a planet's size and its density?
- 25. Consider the fact that the densities of the two rocks that form the majority of Earth's surface, the igneous rocks granite and basalt, are each about 3.0 g/cm<sup>3</sup>. Therefore, the average density of the terrestrial planets is (greater, less than) the density of Earth's surface. Circle your answer.
- **26.** The term (rocky, gaseous) best describes the terrestrial planets. Circle your answer.
- 27. The average density of Earth is about 5.5 g/cm<sup>3</sup>. Considering that the densities of the surface rocks are much less than the average, what does this suggest about the density of Earth's interior?

28.	Which of the planets has a density less than water and therefore would "float"?	36. What is the general relation between the number of moons a planet has compared to its mass? Sug- gest a reason for the relation.
29.	Write a brief statement comparing the densities of the Jovian planets to the density of water.	
	The Jovian planets can be best described as (rocky, ice, and gas) worlds. Circle your answer. Explain why Jupiter can be such a massive object	Rotation and Revolution of the Planets  Rotation is the turning of a planet about its axis that is responsible for day and night. When the solar system is viewed from above the Northern Hemisphere of
32.	Write a general statement comparing the densities of the terrestrial planets to the Jovian planets.	Earth, the planets, with the exception of Venus, rotate in a counterclockwise direction. Venus exhibits a very slow clockwise rotation. The time that it takes for a planet to complete one 360° rotation on its axis is called the <i>period of rotation</i> . The units used to measure a planet's period of rotation are Earth hours and days.  Revolution is the motion of a planet around the
33.	Why are the densities of the terrestrial and Jovian planets so different?	Sun. The time that it takes a planet to complete one revolution about the Sun is the length of its year, called the <i>period of revolution</i> . The units used to measure a planet's period of revolution are Earth days and years. Without exception, the direction of revolution of the planets is counterclockwise around the Sun when the solar system is viewed from above the Northern Hemisphere of Earth.  Use the planetary data in Table 18.1 to answer questions 37–46.
34.	The mass of Pluto, about 0.002 that of Earth's, is most like the masses of the (terrestrial, Jovian) planets, while its density of approximately 2.0 gm/cm <sup>3</sup> is similar to the (terrestrial, Jovian) planets. This suggests that this dwarf planet is made of (solid rock, a rock and ice mixture, all gas). Circle the correct responses.	<ul> <li>37. If you could live on Venus or Jupiter, approximately how long would you have to wait between sunrises?</li> <li>On Venus a sunrise would occur every</li></ul>
The o	aber of Moons of the Planets column labeled "Number of Known Satellites" in the 18.1 indicates the number of known moons or- g each planet.	38. Write a statement comparing the periods of rotation of the terrestrial planets to those of the Jovian planets.
35.	Write a brief statement comparing the number of known moons of the terrestrial planets to the number orbiting the Jovian planets.	The giant planet Jupiter rotates once on its axis approximately every 10 hours. If an object were on the equator of the planet and rotating with it, it would travel approximately 280,000 miles (the equatorial circumference or distance around the equator) in about 10 hours.

39.	Calculate	the	equatorial	rotational	velocity	of
	Jupiter us	ing t	he followin	g formula.		

17-1	Distance _	mi
Velocity $=$ $\frac{1}{2}$	Time	hr
=		mi/hr

40.	The	equatorial	circumference	of	Earth	is	about
	24,00	00 miles. W	hat is the appr	oxi	mate e	qu	atorial
	rotat	ional veloc	ity of Earth?				

miles/	hour
_ 1111103/	Hour

41.	How many times faster is Jupiter's equatorial ro-
	tational velocity than Earth's?

times	faster
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- 42. Compare the planets' periods of rotation to their periods of revolution and then complete the following statement by circling the correct responses. The terrestrial planets all have (long, short) days and (long, short) years, while the Jovian planets all have (long, short) days and (short, long) years.
- **43.** In one Earth year, how many revolutions will the planet Mercury complete and what fraction of a revolution will Neptune accomplish?

Mercury: \_\_\_\_\_ revolutions in one Earth year

Neptune: \_\_\_\_\_ of a revolution in one Earth
year

44. On Venus, how many days (sunrises) would there be in each of its years?

\_\_\_\_\_ day(s) per year

**45.** How many days (rotations) will Mercury complete in one of its years?

Mercury: \_\_\_\_\_ Mercury days in one Mercury year

**46.** Explain the relation between a planet's period of rotation and period of revolution that would cause one side of a planet to face the Sun throughout its year.

In the early 1600s Johannes Kepler set forth three laws of planetary motion. According to Kepler's third law, the period of revolution of a planet, measured in Earth years, is related to its distance from the Sun in astronomical units (one **astronomical unit (AU)** is defined as the average distance from the Sun to Earth—93 million miles or 150 million kilometers). The law states that a planet's orbital period squared is equal to its mean solar distance cubed ( $p^2 = d^3$ ).

47.	Applying Kepler's third law, what would be the
	period of revolution of a hypothetical planet that is 4 AUs from the Sun? Show your calculation in
	the following space.

### **Terrestrial Planet Temperatures**

The temperature of an object is related to the intensity of the heat source, its distance from the source, and the nature of the material it is composed of. To better understand how these variables influence the temperatures of the terrestrial planets, observe the equipment in the laboratory (Figure 18.3) and then complete the following steps. Answer questions 48–51 after you complete your investigation.

- **Step 1:** Working in groups of four or more, obtain four *identical* light (heat) sources and four *identical* black containers with covers and thermometers.
- Step 2: Conduct four experiments simultaneously, one by each member of the group. Do one experiment with the covered can and thermometer 15 cm from the light source, another with the can 30 cm from the light source, the third 45 cm, and the fourth 60 cm.
- **Step 3:** Note the starting temperature for each container on Table 18.3; the temperatures should all be the same.
- **Step 4:** For each of the four setups, turn on the light and record the temperature of the container exactly 10 minutes later. Record the temperatures in Table 18.3.
- **Step 5:** Using the temperature scale on the left axis of the graph, plot the temperatures from Table

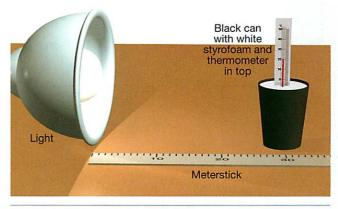


Figure 18.3 Terrestrial planet temperatures lab-equipment setup.

Table 18.3 Temperature Data

Distance from Light Source (cm)	Starting Temperature (°C)	10-minute Temperature
15		
30		
45		
60		

18.3 on the graph in Figure 18.4. Connect the points and label the graph "temperature change with distance."

Step 6: In Table 18.1, notice the mean temperatures for the planets. Plot the mean temperatures of the terrestrial planets at their proper locations on the graph, Figure 18.4. Assume a scale of 40 cm equals 1 AU and use the temperature scale on the right axis of the graph. Label each point with the planet's name. Connect the points and label the graph "mean terrestrial planet temperatures."

48. The "temperature change with distance" graph represents how you would expect that, everything else being equal, the temperature of a planet would be related to its distance from the Sun.

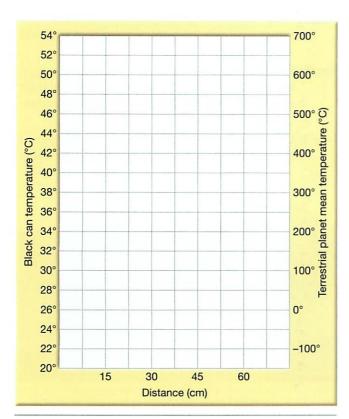


Figure 18.4 Terrestrial planet temperatures graph.

The "mean terrestrial planet temperatures of graph represents the real mean temperatures of the planets. Compare this graph to the theoretica "temperature change with distance graph." How are they different?
for the difference(s) between the two graphs yo
for the difference(s) between the two graphs younged in question 49.
for the difference(s) between the two graphs yo
Write a brief statement suggesting the reason(state of the difference(s) between the two graphs you noted in question 49.  Complete your investigation by writing a statement describing the mean temperatures of the terrestrial planets and the variables that determine those temperatures.
 Complete your investigation by writing a state ment describing the mean temperatures of the terrestrial planets and the variables that determined the complete was a state of the terrestrial planets and the variables that determined the complete was a state of the complete was a sta
for the difference(s) between the two graphs you noted in question 49.  Complete your investigation by writing a statement describing the mean temperatures of the terrestrial planets and the variables that dete

# The Solar System on the Internet

Continue your analyses of the topics presented in this exercise by completing the corresponding online activity on the *Applications & Investigations in Earth Science* website at http://prenhall.com/earthsciencelab

# Patterns in the Solar System

Date Due:	Name:
After you have finished Exercise 18, complete the following questions. You may have to refer to the exercise for assistance or to locate specific answers. Be prepared to submit this summary/report to your instructor at the designated time.  1. On Figure 18.5, prepare a sketch illustrating the planets Mercury, Venus, Earth, and Mars at their approximate distance from the Sun. View the solar system from above the Northern Hemisphere of Earth. Draw arrows around each planet to illustrate its direction of rotation. Also, draw an arrow in the orbit of each planet that shows the direction of revolution.  2. Briefly describe the spacing of the planets in the solar system.	3. Define the following terms:  Terrestrial planets:  Jovian planets:  Plane of the ecliptic:  Rotation:  Mass:  Astronomical unit:

Figure 18.5 Spacing and motion of the terrestrial planets.

(SUN)

0.25 AU

<ol> <li>Referring to the nebular origin of the solar sys- tem, describe and explain the direction of revolu- tion of the planets.</li> </ol>	6. If you knew the distance of a planet from the Sun, explain how you would calculate its period of revolution.
5. Write a brief statement for each of the following characteristics that compares the terrestrial to the Jovian planets.	7. How does a planet's distance from the Sun affect the solar radiation the planet receives? Why?
Diameter:	
Density:	8. How are the mean temperatures of the terrestrial planets related to the solar radiation they intercept? What is the explanation for any discrepancy?
Period of rotation:	
Number of moons:	9. Considering what you have learned about general patterns in the solar system, after reexamining the characteristics of Pluto, discuss why this former planet was reclassified as a dwarf planet in 2006.
Mass:	*